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# Petrogenesis of the High-Alumina Basalt-Andesite suite from Sant'Antioco Island, SW Sardinia, Italy

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ABSTRACT - Present-day Sant'Antioco Island (SW-Sardinia, Italy) hosted one of the southernmost, and youngest, subduction-related, Cenozoic magmatic events of western Sardinia. A high-alumina basaltandesite rock association, the focus of this paper, crops out in the southern portion of the island and represents the mafic end-member of volcanism developed during Miocene. Basaltic andesites and andesites are the dominant rock-types, while basalts and dacites are occasional. Minero-petrographic and geochemical characteristics reveal: i) a clear calc-alkaline signature for these rocks, in spite of high FeO\*/MgO ratios that mimic a tholeiitic affinity; ii) a magma evolution mainly controlled by fractional crystallization and *iii*) some degree of crustal assimilation. Phase relationships and compositions, mass balance calculations and T, P, X<sub>HaO</sub> estimates consistently indicate that fractional crystallization occurred at low pressure (P=100-400 MPa), under different  $P_{H_{2}O}$ conditions, which explain the observed wide shifts in major oxides (e.g., Al<sub>2</sub>O<sub>3</sub> and MgO), for comparable SiO<sub>2</sub> contents, and the ambiguous tholeiitic character as well. Phase relationships in the least evolved lithotypes point out that higher H<sub>2</sub>O concentrations (estimated at up to 6 wt%) in the melts reduced the crystallization of plagioclase and favoured that of olivine, whereas lower H<sub>2</sub>O contents (up to 3 wt%) promoted plagioclase fractionation and earlier crystallization of orthopyroxene at the expense of olivine. This X<sub>H-O</sub> imprint was recorded up to the most evolved compositions. Finally, the study rock suite can be derived from a parental (near-primary) magma matching in composition high-MgO basalts from the Montresta calc-alkaline district in NW Sardinia, in the light of the close compositional similarity between the high-Al<sub>2</sub>O<sub>3</sub> basaltic andesites in the two districts. Thus, starting from a common parental composition, here we reconstruct liquid lines of descent, where the relative proportions of fractionating plagioclase vs mafic phases, as controlled by  $X_{H_{2}O}$  in the melts, may account for the compositional differences of Sant'Antioco rock-types ab initio.

RIASSUNTO - L'Isola di Sant'Antioco (Sardegna sudoccidentale) ospita una delle più recenti manifestazioni del magmatismo orogenico del margine occidentale della Sardegna connesso alla subduzione cenozoica. Oggetto del presente lavoro sono le vulcaniti affioranti nella porzione meridionale dell'isola, costituite essenzialmente da andesiti basaltiche alte in allumina e da andesiti (più occasionali basalti e daciti), che rappresentano i termini meno differenziati del locale magmatismo di età miocenica. Oltre ad integrare l'ampio

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panorama degli studi esistenti sul magmatismo orogenico sardo, la caratterizzazione minero-petrografica e geochimica e la parametrizzazione P, T, X<sub>H2O</sub> dei litotipi investigati, consente una ricostruzione dei relativi processi petrogenetici, anche alla luce dei recenti sviluppi sperimentali sui sistemi calco-alcalini. I risultati indicano che le vulcaniti in esame rivelano una chiara affinità calco-alcalina, nonostante gli alti rapporti FeO\*/MgO tipici di rocce di serie tholeiitica, e che la differenziazione magmatica è stata controllata principalmente dalla cristallizzazione frazionata, associata a differenti gradi di assimilazione crostale. Le relazioni di fase e i calcoli di bilancio di massa, in accordo con le stime di T e P di cristallizzazione dei fenocristalli e dei contenuti di H2O in soluzione nei magmi, suggeriscono una cristallizzazione a bassa pressione (P=100-400 MPa) con differenti X<sub>H<sub>2</sub>O</sub>. Proprio quest'ultimo parametro spiega sia l'ampia variabilità composizionale di queste rocce (in particolare i differenti contenuti in Al<sub>2</sub>O<sub>3</sub> e MgO per contenuti in SiO<sub>2</sub> simili), sia l'apparente carattere tholeiitico. Le relazioni di fase osservate nei litotipi meno evoluti indicano infatti che alte concentrazioni di H2O nei fusi (stimate intorno al 6% in peso) hanno ridotto l'incidenza del plagioclasio (che tuttavia rappresenta invariabilmente la fase dominante, seguita dal clinopirosseno), favorendo la cristallizzazione dell'olivina fra le fasi femiche. Viceversa. minori contenuti in  $H_2O$  (fino a ~3%), oltre a favorire il frazionamento di plagioclasio, hanno comportato la precoce cristallizzazione dell'ortopirosseno a spese dell'olivina. L'evidenza di diversi contenuti in H<sub>2</sub>O quale fattore di controllo determinante del processo di differenziazione dei magmi persiste fino alle composizioni più evolute dell'associazione in esame. Infine, le litologie meno differenziate riscontrate a Sant'Antioco possono essere considerate come direttamente derivanti da un magma più primitivo di composizione simile a quella dei basalti alti in MgO affioranti nel distretto vulcanico di Montresta (Sardegna nord-occidentale), anche sulla base delle significative analogie composizionali fra le andesiti basaltiche alte in Al<sub>2</sub>O<sub>3</sub> nelle due aree. L'insieme dei dati è compatibile con un modello in cui, a partire da una comune composizione parentale, le differenze riscontrate sin dai termini meno evoluti di Sant'Antioco possono essere ricondotte a percorsi evolutivi caratterizzati da diversi rapporti tra le fasi frazionate (in particolare fra il plagioclasio e le fasi femiche) sotto il controllo cruciale della X<sub>H2O</sub>.

KEY WORDS: Cenozoic orogenic volcanism, highalumina basaltic-andesites, petrology and geochemistry, Sant'Antioco, Sardinia.

#### INTRODUCTION

Cenozoic, subduction-related, igneous activity in Sardinia has been widely investigated (e.g., Dostal et al., 1982; Beccaluva et al., 1994; Brotzu et al., 1997a, b; Morra et al., 1994, 1997; Mattioli et al., 2000; Downes et al., 2001; Franciosi et al., 2003; Lustrino et al., 2004, 2009). However, information on petrography, mineralogy and chemistry of one of the southernmost volcanic districts, Sant'Antioco, is still scarce and incomplete. Up to now, the few geological and compositional data available on the Sant'Antioco volcanics were reported in Luxoro (1987), in the 1:25,000 geo-petrographic map of Sant'Antioco Island (Maccioni et al., 1990) and in Lecca et al. (1997), who considered the andesitic s.l. rocks from Sant'Antioco Island as part of the orogenic magmatism in the Sulcis area (SW mainland Sardinia).

The Sulcis volcanic district is the southernmost part of the calc-alkaline magmatic belt that developed during late Eocene - middle Miocene time span (Lustrino *et al.*, 2009) along the western side of Sardinia (Fig. 1), as a consequence of subduction acting in the Western Mediterranean Sea beneath the southern European margin (e.g., Doglioni, 1991; Doglioni *et al.*, 1999 and references therein).

This work illustrates the mineralogical, geochemical and petrological aspects of the Sant'Antioco volcanics in order to integrate the large set of studies on Cenozoic orogenic magmatism in Sardinia and to reconstruct petrogenetic processes. In particular, we focus on the influence of H<sub>2</sub>O on the magma liquid lines of descent in the light of recent studies on natural and experimental calc-alkaline systems (Sisson and Grove, 1993a, b; Pichavant and Macdonald, 2007 and references therein; Ulmer, 2007 and



Fig. 1 - a) Geological sketch map of Sardinia, showing the location of Sant'Antioco Island and other subduction-related Cenozoic volcanics (after Cherchi and Montadert, 1982, Savelli *et al.*, 1979 and Morra *et al.*, 1997, modified). (1) Pliocene-Quaternary sediments; (2) Middle Miocene-Pleistocene anorogenic volcanics; (3) Late Eocene-Middle Miocene sedimentary and volcanic rocks: (3a) silicic ignimbrites; (3b) andesites *s.l.*; (4) continental sedimentary rocks of Cixerri Formation; (5) undifferentiated Paleozoic basement and Mesozoic up to Eocenic cover; (6) main Post-Paleozoic regional faults. b) Geological sketch map of Sant'Antioco Island. c) Enlargement of the area shown in b, where the sampling sites are reported. (\* lava clasts in Pispisia-M.te Arbus polygenic breccias).

references therein). Furthermore, possible relationships with compositionally analogous, or even less differentiated, Cenozoic orogenic rocktypes of western Sardinia will be considered, aiming at identifying a reliable near-primary parental magma for the studied rock association.

#### GEOLOGICAL BACKGROUND

The Sant'Antioco Island, located immediately offshore of southwestern mainland Sardinia (Italy; Fig. 1), extends 17.5 km in a N-S direction, with a maximum W-E width of 10 km. It is part of the calc-alkaline volcanic belt that stretches along the western coast of Sardinia, which developed during late Eocene-middle Miocene (38-12 Ma, Lustrino et al., 2009 and references therein) along the main branches of the "Sardinian Rift System" (Main Branch Volcanic Zone) and in Sulcis (South-Western Volcanic Zone; sensu Lecca et al., 1997). The Sardinian Rift System and related magmatism are interpreted in the frame of the kinematic processes acting in the western Mediterranean, which involved the Hercvnian paleo-margin of southern Europe. These processes occurred in connection with the back-arc spreading produced by the eastward roll-back of the westwardsubducting plate related to convergence between Africa and Europe (Doglioni, 1991; Doglioni et al., 1999 and references therein). This caused the separation counterclockwise and (i.e., southeastward) rotation of the Corsica-Sardinia continental microplate up to its present position (upper Oligocene-middle Miocene). In this context, the Cenozoic calc-alkaline magmatism of Sardinia is directly related to the subduction of oceanic crust. The earliest calc-alkaline magmatic events along the present western side of Sardinia predate the rotation of the Corsica-Sardinia microplate; then magmatism migrated southeastward following the west-directed, eastward retreating, subduction zone beneath Sardinia until middle Miocene. The Cenozoic

orogenic magmas in Sardinia were emplaced through a variably thinned continental crust (ca. 30 to 20 km thick, from North to South; Egger *et al.*, 1988; Sartori *et al.*, 2004; Sau *et al.*, 2005), which may account for the differences in composition and eruptive style of magmatic activity from North to South along the Sardinian Rift (e.g., Lustrino *et al.*, 2002).

The Sulcis volcanic district, which includes a south-western coastal area of mainland Sardinia and the nearby islands of San Pietro and Sant'Antioco, represents the southernmost expression of orogenic magmatism in Sardinia (Fig. 1) and one of its youngest occurrences as well. Volcanic stratigraphy is characterized by: 1) lava and subordinate pyroclastic successions, including andesites, basaltic andesites and rare basalts, of Oligo-Miocene age, which crop out in Sulcis mainland (i.e., S. Giovanni Suergiu, Narcao, Giba) and in the southern part of Sant'Antioco, and 2) silicic welded ignimbrites and subordinate lava flows and domes, ranging in composition from dacites to comendites, of Miocene age, exposed in mainland Sulcis (Portoscuso area), in the northern and western parts of Sant'Antioco, and throughout the whole San Pietro Island (Assorgia et al., 1990; Morra et al., 1994; Pioli and Rosi, 2005; Pioli et al., 2008).

The pre-volcanic substrate on Sant'Antioco includes biohermal-biostromal and oolitic limestones and silty marls of Cretaceous age, cropping out in the SE part of the island (Maccioni et al., 1990). Almost all of the island is made up of volcanic rocks of Miocene age from both silicic, mainly explosive, and mafic to intermediate, mainly effusive, activities. Published K-Ar age data indicate values clustered at around 16-18 Ma (Araña et al., 1974; Savelli et al., 1979; Montigny et al., 1981; Maccioni et al., 1991). A younger age of about 12 Ma was recently obtained by40Ar/39Ar determinations for the trachytic neck of Isola del Toro, South of Sant'Antioco (Lustrino et al., 2007), which is intermediate between the age of the last products

of the Oligo-Miocene orogenic igneous phase in SW Sardinia and the early volcanism of the late Miocene-Quaternary phase in SE Sardinia (ca. 6.64 Ma, Lustrino *et al.*, 2007).

# FIELD CHARACTERISTICS OF THE STUDIED ROCK-TYPES

The present study investigates the mafic to intermediate volcanic rocks that crop out in the southern part of Sant'Antioco Island. These are represented by high- and low-aspect ratio lava flows and small lava domes (typically less than 1 km across), associated with minor pyroclastic deposits, locally cut by lava dykes. In order to characterize this volcanic activity on Sant'Antioco, a set of sixty-seven representative rock samples has been analyzed, as listed in TABLE 1 (sample localities are shown in Fig. 1). TABLE 1 also summarizes relevant field characteristics of the studied volcanics, along with their correlation with the units identified by previous studies (i.e., Luxoro, 1987; Maccioni et al., 1990). Following the stratigraphic scheme of Luxoro (1987), the studied samples include: i) twenty-two samples from the "Lave iniziali" (early lavas, hereon abbreviated as LI), consisting of thick (up to 10 m), massive, lava

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List of the analyzed samples of Cenozoic mafic to intermediate volcanic rocks from Sant'Antioco Island studied in this work (sample locations in Fig. 1), grouped according to unit nomenclatures in Luxoro (1987) and Maccioni et al. (1990). L1: Lave iniziali (Early lavas); P: Pispisia-M.te Arbus polygenic volcanic breccias; LT: Lave terminali (Final lavas); F: lava dykes.

SAMPLES	FIELD ASPECT	Luxoro (1987)	Maccioni et al. (1990)
SA22, SA29-29b, SAV8b	Lava dyke	F	13a
SA21, SA24, SA25	Lava dyke	F	13b
SA1, SA2, SA3-3b, SA4, SA6, SA7, SA8, SA9, SA16, SA35, SA36, SA37, SA38, SA39, SA40-40b, SAV2, SAV4, SAV11, SAV12	Thin lava flow interbedded with scoriaceous levels	LT	14
SAV16, SAV17, SAV19, SAV21	Lava clasts in polygenic breccias	Р	15a
SA19, SA27, SA28, SAV14, SAV15-15b, SAV20	Lava injections in polygenic breccia:	s P	15b
SA15-15b, SA23, SAV18	Thick lava flow or lava dome	LI	17a
SA5, SA10, SA11, SA30, SA34-34b, SAV7, SAV9, SAV10, SAV13	Thick lava flow	LI	18
SA12-12b, SA13, SA14, SA17, SA18, SA31, SA33, SAV5, SAV6, SAV8	Thick lava flow or lava dome	LI	19a
SA20, SA32-32b	Lava dome	LI	20

flows and lava domes; *ii*) eleven samples from lava injections and lava blocks within the Pispisia-M.te Arbus polygenic volcanic breccia deposits (hereon P); *iii*) twenty-seven samples from thin (usually 1.5 m-thick) lava flows, part of the "Lave terminali" (final lavas, hereon LT); *iv*) seven samples from lava dykes intersecting the above volcanics (hereon F).

#### Petrography

The studied Sant'Antioco volcanics range in composition from basalts to dacites (see Geochemistry section). Basaltic andesites largely prevail within LT rock-types, while andesites are dominant within LI lithologies. Both basaltic andesites and andesites occur in the Pispisia-M.te Arbus lava types (P), as well as in lava dykes (F). Hereafter, for the sake of simplicity, the products from P and F units are grouped within either the LI or LT main-types, on the basis of bulk rock composition and of common mafic phenocryst assemblages (plagioclase is invariably the main phase and opaques are ubiquitous, see below). Lava rocks from both LI and LT mostly show porphyritic or glomeroporphyritic (15-48 vol% phenocrysts), isotropic to pilotaxitic or seriate, textures and microcrystalline groundmass. Phenocryst and groundmass assemblages, as determined from modal analyses in thin section, are reported in TABLE 2 for a set of representative samples.

LT rocks (also including part of F dykes) consist of basaltic andesites and occasional basalts characterized by the ubiquitous, although relatively scarce, presence of olivine (always <5 vol%) and by the prevalence of clinopyroxene (Cpx) on orthopyroxene (Opx). In a few cases, within the least evolved lavas, which show very low contents of mafic phenocrysts (about 5%), Opx is lacking and olivine (3-4 vol%) prevails on Cpx.

LI rocks (also including P lavas and part of F dykes) comprise basaltic andesites and occasional dacites characterized by: *i*) the ubiquitous presence of pyroxene phenocrysts, with Opx often prevailing on Cpx; *ii*) the lack or occasional presence of scarce amounts of olivine; *iii*) the occurrence of hornblende in two samples from the

Rock-type	LT	LT	LT	LT	LI	LI	LI	LI
Sample:	SA7 HABA	SA40 HABA	SA35 HABA	SA3 HABA	SA12 AND	SA15 AND	SA31 AND	SA32 DAC
Plg (ph <sup>a</sup> +mph <sup>b</sup> )	23.2	23.5	16.8	21.9	22.5	21.4	32.1	29.5
Ol (ph+mph)	3.5	5.4	3.5	0.2	-	-	-	-
Cpx (ph+mph)	6.6	4.2	0.6	9.1	0.6	2.5	3.4	2.7
Opx (ph+mph)	-	-	-	-	1.6	7.6	9.7	6.5
Opx mph	0.5	-	-	0.8	-	-	-	-
Fe-Ti oxides mph	1.0	1.1	0.9	1.5	1.3	2.3	2.5	1.9
Hbl ph	-	-	-	-	-	-	-	1.0
P.I. (ph+mph)	34.8	34.2	21.8	33.5	26.0	33.8	47.7	41.6
Gdm	65.2	66.8	78.2	66.5	74.0	66.2	52.3	58.4

TABLE 2

Modal analyses of the most representative LT and LI rock-types (at least 1000 points counted for each sample).

<sup>a</sup> phenocrysts (> 0.5 mm); <sup>b</sup> microphenocrysts (> 0.05 mm); Plg: plagioclase; OI: olivine; Cpx:

clinopyroxene; Opx: orthopyroxene; Hbl: horneblende; P.I.: porphyritic index; Gdm: groundmass; HABA: high-Al<sub>2</sub>O<sub>3</sub> basaltic andesite; AND: andesite; DAC: dacite.

most evolved andesites and in the dacite sample.

Phenocryst glomeri (mostly Cpx and Opx, associated with coarse-grained Fe-Ti oxides, subordinate plagioclase and occasional olivine) are more common in LI than LT and typically occur in two-pyroxene andesites (cfr. "Type a" microinclusions in Brotzu *et al.*, 1997a).

Mineral chemistry and geochemical features of whole rock samples are reported in the following sections. Analytical methods are illustrated in Appendix A.

#### MINERALOGY

#### Plagioclase

Plagioclase is the most abundant phenocryst phase (up to 80% by volume of total phenocrysts and microphenocrysts) in both LT and LI groups. It usually occurs as individual, variably sized, euhedral grains or, mainly in andesites, in glomerocrystic aggregates. Plagioclase crystals often show continuous zoning and clear surfaces. Dusty or cellular textures may be present, where crystal portions may enclose glassy material. Plagioclase phenocrysts also enclose magnetite and minute plagioclase crystals or, more rarely, fine skeletal Cpx grains.

In LI rock-types, plagioclase crystals show common intracrystalline compositional ranges, with either normal (core: An<sub>75-67</sub>; rim: An<sub>65-59</sub>) or reverse (core: An<sub>64-68</sub>; rim: An<sub>72-74</sub>) zoning. Although usually normally zoned. LT plagioclases are characterized by a wider compositional pattern: phenocryst core compositions even  $>An_{90}$  are observed in several examples. while phenocryst rim and microphenocryst mostly range  $An_{60-75}$ in (representative composition analyses are reported in TABLE 3). As a whole, the compositional ranges of LT plagioclase vary significantly, even within rocks with similar evolution degree. For example, if we consider the least evolved lithotypes, the prevailing plagioclase compositions vary from  $An_{75-85}$  (SA7), through  $An_{78-90}$  (SA40), to  $An_{85-93}$  (SA35) (TABLE 3; Fig. 2), concomitant with increasing  $Al_2O_3$  contents in the corresponding whole rocks (cfr. Geochemistry section).

#### Pyroxenes

In LT rocks, Cpx is the dominant mafic phenocryst phase (5-15% by volume), while Opx usually does not exceed 5 vol%. With increasing degree of differentiation, the amount of Opx vs Cpx increases to become prevalent in most of the olivine-free LI samples. In these latter, Cpx and Opx phenocrysts are often associated in polycrystalline aggregates (glomeri). As a whole, Cpx may occur as: i) euhedral pheno- to microphenocryst grains; ii) coarser, rounded grains, sometimes enclosing corroded relics of olivine; iii) microgranular coronas that mantle both Opx and Cpx phenocrysts, mainly in evolved LI andesites; iv) microliths in the groundmass. Clinopyroxene is always augitic, with a narrow compositional range (Wo37-45 En39- $_{45}$  Fs<sub>11-20</sub>). No clear compositional variations are observed as a function of the degree of evolution from basalts to andesites: both Al<sub>2</sub>O<sub>3</sub> (1.6-6.5 wt%) and TiO<sub>2</sub> (0.33-1.10 wt%) contents do not well correlate with the magnesium number [mg# = MgO/(MgO+FeO\*) molar], as they show different amounts for comparable mg# values (e.g., LT in Fig. 3 and TABLE 4).

Orthopyroxene ( $Wo_{2-4} En_{54-72} Fs_{25-42}$ , TABLE 4) occurs as either euhedral to subhedral phenocrysts or microphenocrysts (either isolated or in polycrystalline glomeri), in some cases mantled by Cpx grains. Solid inclusions, commonly of magnetite and, more rarely, of plagioclase laths, are also found.

### Olivine

Olivine mostly occurs in LT rocks and usually does not exceed 2-3 vol%. In a few, less evolved,

Unit	SA7 LT	SA7 LT	SA7 LT	SA7 LT	SA7 LT	SA7 LT	SA35 LT	SA35 LT	SA35 LT	SA35 LT	SA35 LT	SA35 LT	SA40 LT	SA40 LT	SA40 LT	SA40 LT	SA40 LT
	1/c	1/r	2/c	2/r	mph	аз	1/c	1/r	2/c	3/c	nph	аз	1/c	2/c	2/r	hqm	аз
$SiO_2$ Al,O <sub>3</sub>	45.45 34.88	48.78 31.26	47.43 33.05	47.54 32.92	47.73 32.38	48.99 31.65	44.89 34.05	49.64 31.19	43.61 35.92	45.89 34.68	46.44 33.30	49.21 31.92	44.42 35.55	45.50 31.74	47.65 32.97	46.82 33.71	47.86 32.59
FeO	0.84	0.63	0.85	0.64	0.90	0.76	0.54	0.97	0.62	0.64	0.81	0.93	0.58	0.66	0.61	0.67	0.83
CaO	17.77	16.18	16.62	15.85	15.70	14.97	19.70	15.73	19.25	17.87	16.91	14.99	18.68	17.71	15.80	16.83	15.57
$Na_2O$	1.56	2.88	2.19	2.73	2.77	3.27	0.87	3.03	0.74	1.53	2.18	3.25	1.14	1.71	2.87	2.21	2.86
$ m K_2O$	ı	0.09	0.10	0.10	0.10	0.10	0.04	0.13	ı	ı	·	0.12	ı	ı	ı	ı	0.18
Total	100.49	99.82	100.24	99.78	99.58	99.74	100.09	100.69	100.14	00.62	99.64	100.41	00.37	97.32	99.90	00.24	99.89
An (mol %)	86.32	75.26	80.28	75.80	75.37	71.26	92.39	73.62	93.50	86.57	81.08	71.35	90.05	85.13	75.26	80.80	74.28
Ab (mol %)	13.68	24.24	19.14	23.63	24.06	28.17	7.38	25.66	6.50	13.43	18.92	27.99	9.95	14.87	24.74	19.20	24.69
Or (mol %)	0.00	0.50	0.58	0.57	0.57	0.57	0.22	0.72	0.00	0.00	0.00	0.66	0.00	0.00	0.00	0.00	1.02
$\mathrm{K}_\mathrm{D}$	2.40	1.18	1.59	1.22	1.19	0.96	5.12	1.17	5.88	2.63	1.75	1.04	4.12	2.60	1.38	1.91	1.37
	SAV1	SAV11	SAV11	SAV11	SAV11	SAV11	SAV15	SAV15	SAV15	SAV15	SAV15	SAV5	SAV5	SAV5	SAV5	SAV5	SAV5
Unit	LT	LT	LT	LT	LT	LT	Ь	Ь	Ь	Ь	Ь	ΓI	ΓI	LI	LI	ΓI	ΓI
	1/c	1/r	2/c	2/r	nph	а	1/c	1/r	2/c	mph	а	1/c	1/r	2/c	2/r	mph	а
$SiO_2$	46.54	50.87	45.18	45.32	53.86	55.24	46.15	47.36	47.42	48.70	59.11	49.32	53.02	50.32	52.43	51.31	59.31
${ m Al}_2 { m  ilde O}_3$	34.12	30.80	34.73	34.53	28.42	27.33	34.23	33.12	33.33	32.38	23.92	32.22	29.62	31.40	29.94	30.51	24.36
FeO	0.42	0.74	0.56	0.81	1.28	1.51	0.72	0.61	0.69	0.78	1.03	0.33	0.44	0.52	0.30	0.64	0.86
CaO	17.07	13.51	18.14	17.98	10.88	9.66	17.25	16.53	16.41	15.12	5.94	15.15	12.10	14.13	12.47	13.19	6.26
$Na_2O$	1.72	3.79	1.06	1.28	5.09	5.45	1.54	1.97	2.04	2.72	5.53	2.84	4.49	3.40	4.31	3.97	5.31
$ m K_2O$		ı			0.29	0.72	•	ı	0.13	ı	3.66	ı	0.22	0.13	0.20	ı	3.88
Total	99.87	99.71	99.67	99.92	99.82	99.91	99.89	99.59	100.02	99.70	99.19	99.86	99.89	99.90	99.65	99.62	99.98
An (mol %)	84.58	66.33	90.44	88.59	53.24	47.40	86.09	82.26	81.01	75.44	29.25	74.67	59.06	69.14	60.81	64.74	30.55
Ab (mol %)	15.42	33.67	9.56	11.41	45.07	48.39	13.91	17.74	18.22	24.56	49.28	25.33	39.66	30.10	38.03	35.26	46.90
Or (mol %)	0.00	0.00	0.00	0.00	1.69	4.21	0.00	0.00	0.76	0.00	21.46	0.00	1.28	0.76	1.16	0.00	22.55
$\mathbf{K}_{\mathrm{D}}$	2.42	0.87	4.16	3.42	0.52	0.43	2.96	2.22	2.13	1.47	0.28	1.73	0.87	1.35	0.94	1.08	0.38
Note: c: pher exchange dis	nocryst c tributior	ore; r: p 1 coeffic:	henocry: ients (K <sub>I</sub>	st rim (1 ) are als	, 2, 3: ph so report	ienocrys ed (see 1	st identif text for e	ication) calculati	; mph: n on detai	nicroph∈ ls).	enocrys	t; g: gro	undmas	s. The p	lagiocla	se-melt	Ca-Na

TABLE 3 Representative electron microprobe analyses of plagioclase from the Sant'Antioco study rock-types.

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Fig. 2 - Frequency histograms displaying the variation of An contents in plagioclase of the SA7, SA35 and SA40 highalumina basaltic andesite (HABA) samples from LT unit, which stand for the least evolved Sant'Antioco products (frequency expressed as number of analyses). Note that each sample shows a dominant plagioclase composition related to the Al<sub>2</sub>O<sub>3</sub> content in the whole rock (cfr. TABLE 6).

rocks it may attain up to 5 vol% and represent the main mafic phase. Overall, Fo component ranges 50-59 mol%; individual crystals are usually normally zoned (TABLE 5). Generally, olivine crystals are altered to iddingsite, even if they preserve the original habit. In a few cases, fresh olivine relics are observed at the core of altered crystals. Some olivine crystals are bordered by granular Cpx.

# Amphibole

Euhedral to subhedral amphibole phenocrysts, showing opacitic rims, are found in scarce amounts (<2 vol%) in two LI andesite samples with SiO<sub>2</sub> contents >58 wt% and in the dacite sample. Representative analyses, reported in TABLE 5, allow a classification as MgO-hornblende, following Leake *et al.* (2004). Compositional (i.e., low Al<sub>2</sub>O<sub>3</sub> content, ~8 wt% on average) and textural relationships testify their late stage crystallization under low pressure conditions.

# Fe-Ti oxides

Fe-Ti oxides (TABLE 5) are present ubiquitously both as microphenocrysts (either dispersed or bordering pyroxene phenocrysts) and, in higher amounts, as groundmass phases. Magnetite-titanomagnetite is common in basaltic andesites, while ilmenite only occurs in andesites.

# GEOCHEMISTRY

Representative bulk rock compositions from the Sant'Antioco volcanics are reported in TABLE 6. Essentially they plot in the basaltic andesite and andesite fields of the TAS classification diagram (Le Bas et al., 1986), with quite subordinate basalts and dacites. According to other classification diagrams, i.e. K<sub>2</sub>O vs SiO<sub>2</sub> (Peccerillo and Taylor, 1976; Fig. 4a), and MgO vs Al<sub>2</sub>O<sub>3</sub> (Kersting and Arculus, 1994; Fig. 4b), the studied products are mainly high-alumina basaltic andesites and andesites (with quite subordinate basalts and dacites: only one data point each) of medium-K calc-alkaline affinity. On the other hand, on the mere basis of the FeO\*/MgO vs SiO<sub>2</sub> diagram (Miyashiro, 1974) all the investigated rocks would show a tholeiitic to mildly tholeiitic affinity (Fig. 4c). However, the plot of FeO\*/MgO vs FeO\* (inset in Fig. 4c) indicates that the relatively high values of FeO\*/MgO in the Sant'Antioco rock suite are not accompanied by an iron enrichment trend typical of tholeiitic series. Furthermore, a possible tholeiitic affinity for the Sant'Antioco volcanics conflicts with the lack of pigeonite microphenocrysts (Kuno, 1959) and of Feenrichment in plagioclase (Aramaki and Katsura, 1973), as well as with the observed  $TiO_2$  and Al<sub>2</sub>O<sub>3</sub> contents, which pertain more properly to calc-alkaline rock-series. On these grounds, the relatively high FeO\*/MgO ratios, which mimic a tholeiitic affinity, may be likely linked to the relative proportions of fractionating plagioclase

	Representa	ative elec.	tron micro	pprobe and	alyses of	orthopyro	xenes (a)	and clin	юругохеп	es (b) from	the Sant	Antioco s	tudy rock-t	vpes.	
(a) Unit	SAV12 LT r	c SA7 LT c	SAV11 LT c	SAV11 LT r	SAV15 P c	SAV15 P r	SAV15 P mph	SAV15 P g	SAV13 LI r	SAV13 LI mph	SA10 LI c	SAV5 LI c	SAV5 LI r	SAV8 LI c	SAV8 LI r
SiO <sub>2</sub> TiO <sub>2</sub> Al <sub>2</sub> O <sub>3</sub> FeO MnO MgO Na <sub>2</sub> O <i>Total</i>	52.18 52.18 0.25 19.99 0.35 1.51 0.27 99.66	53.77 53.77 0.33 1.05 16.99 0.48 0.48 0.48 1.82 1.82	$\begin{array}{c} 51.25\\ 0.35\\ 0.35\\ 0.35\\ 18.73\\ 0.40\\ 0.40\\ 24.61\\ 1.54\\ 0.19\\ 0.19\\ 99.69\end{array}$	53.55 0.27 1.23 18.39 0.62 25.20 1.50 1.50	52.48 0.31 1.39 19.58 0.54 24.12 1.61 0.18 0.18	51.49 0.42 2.24 19.77 0.35 0.35 1.75 1.75 100.05	$\begin{array}{c} 52.80\\ 0.30\\ 0.30\\ 2.10\\ 17.45\\ 0.34\\ 0.34\\ 1.59\\ 0.46\\ 0.46\\ 0.46\end{array}$	$\begin{array}{c} 52.72\\ 52.72\\ 0.25\\ 1.77\\ 19.99\\ 0.35\\ 0.35\\ 0.28\\ 0.28\\ 0.28\end{array}$	51.40 0.29 0.38 0.33 0.93 19.21 1.72 1.72	52.19 0.38 0.44 24.71 1.00 19.81 1.88 1.88 1.00.41	$\begin{array}{c} 52.19\\ 0.21\\ 0.21\\ 0.83\\ 0.69\\ 0.69\\ 1.77\\ 1.77\\ 0.03\\ 100.24\end{array}$	51.68 0.28 1.36 24.51 0.88 19.96 1.17 0.18 0.18	$\begin{array}{c} 51.73\\ 0.16\\ 1.37\\ 24.86\\ 0.75\\ 19.72\\ 1.47\\ 1.47\\ 0.34\\ 100.40\end{array}$	52.16 0.38 0.62 0.62 23.58 23.58 0.78 0.78 1.86 1.86 99.81	52.17 0.30 0.39 0.39 24.41 0.93 0.93 20.18 1.64 1.64
Wo (mol%) En (mol%) Fs (mol%) mg#	3.02 65.55 31.43 0.68	3.56 69.79 26.65 0.73	3.05 67.92 29.03 0.70	2.93 68.39 28.69 0.71	3.19 66.56 30.25 0.69	3.46 66.13 30.40 0.68	3.20 68.89 27.92 0.72	3.03 65.53 31.44 0.68	3.52 54.55 41.94 0.56	3.82 56.00 40.18 0.59	3.52 63.52 32.95 0.65	2.40 56.95 40.65 0.59	3.01 56.11 40.89 0.59	3.78 57.78 38.44 0.61	3.33 57.06 39.60 0.60
(b) Unit	SAV12 LT c	SA7 LT c	SA35 LT c	SAV15 P c	SAV15 P r	SAV15 P mph	SAV10 LI c	SA15 LI c	SA15 LI r	SA15 LI mph	SA15 LI r	SA10 LI c	SA10 LI r	SAV8 LI c	SAV8 LI r
SiO <sub>2</sub> TiO <sub>2</sub> Al <sub>2</sub> O <sub>3</sub> FeO MnO MgO CaO CaO CaO Cr <sub>2</sub> O <sub>3</sub> <i>Total</i>	50.77 0.57 2.67 2.67 10.44 0.16 15.27 19.08 0.39	$\begin{array}{c} 51.01\\ 0.52\\ 0.52\\ 11.86\\ 11.42\\ 0.32\\ 16.06\\ 18.36\\ 0.24\\ 99.79\end{array}$	48.03 1.09 5.12 9.74 0.30 0.30 0.48 0.48	48.57 0.88 6.32 6.32 7.34 0.15 0.15 0.55 99.11	49.57 0.85 5.79 7.88 7.88 14.78 20.53 0.44 100.04	$\begin{array}{c} 50.63\\ 0.79\\ 0.79\\ 10.67\\ 0.46\\ 15.11\\ 18.45\\ 0.27\\ 0.13\\ 99.93\end{array}$	$\begin{array}{c} 49.66\\ 0.67\\ 3.96\\ 111.77\\ 0.39\\ 0.39\\ 1.4.11\\ 18.46\\ 0.52\\ 0.52\\ 99.54\end{array}$	$\begin{array}{c} 49.66\\ 0.67\\ 3.96\\ 111.77\\ 0.39\\ 0.39\\ 0.52\\ 0.52\\ 99.54\end{array}$	50.44 0.79 0.79 3.88 10.45 0.23 13.98 19.45 0.40 0.40	51.51 0.33 2.55 10.05 0.00 14.95 19.68 0.36 0.36	50.44 0.79 0.79 3.88 10.45 0.23 13.98 19.45 0.40 0.40	$\begin{array}{c} 49.63\\ 0.70\\ 3.13\\ 11.30\\ 0.30\\ 14.27\\ 20.36\\ 0.34\end{array}$	50.91 0.33 1.64 12.94 0.45 14.14 19.65 0.31 0.31	51.51 0.33 2.55 10.05 0.00 14.95 19.68 0.36 0.36	50.59 0.58 0.58 11.94 0.23 13.55 19.98 0.25 0.25
Wo (mol%) En (mol%) Fs (mol%) mg#	39.26 43.72 17.03 0.72	37.19 45.27 17.54 0.71	43.08 40.66 16.27 0.72	44.75 42.75 12.50 0.78	43.31 43.39 13.31 0.77	38.31 43.65 18.04 0.72	38.79 41.26 19.95 0.68	38.79 41.26 19.95 0.68	41.17 41.18 17.65 0.70	40.72 43.05 16.23 0.73	41.17 41.18 17.65 0.70	41.58 40.55 17.87 0.69	39.98 40.03 20.00 0.66	40.72 43.05 16.23 0.73	41.33 39.01 19.66 0.67
c: phenocrys	t core ; r: J	henocry:	st rim; m	ph: micro	phenocr	yst; g: gr	oundmas	ss; mg#:	MgO/(N	IgO+FeO)	) molar, w	vith all irc	in as Fe <sup>2+</sup> .		

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Fig. 3 - Variation diagrams of magnesium number (mg#) vs  $Al_2O_3$  (a) and  $TiO_2$  (b) for clinopyroxenes of representative Sant'Antioco samples. Note the difference in  $Al_2O_3$  and  $TiO_2$  contents of clinopyroxenes in lavas with similar evolution degree (i.e., SA7 and SA35 considered for petrogenetic modeling). HAB: high-alumina basalt; HABA: high-alumina basaltic andesite; AND: andesite.

and mafic phases during magma evolution, leading to the observed low-MgO and high-Al<sub>2</sub>O<sub>3</sub> compositions (e.g., Pichavant and McDonald, 2007; also see below).

Variation diagrams (Fig. 5) exhibit broad correlations of major oxides with MgO, besides wide scatter (e.g., Al<sub>2</sub>O<sub>3</sub> and CaO). As a whole, in spite of distinct field appearance and petrographic features, LI and LT do not define separate compositional trends, whereas both rock groups display correlations consistent with crystallization and removal of the aforementioned phenocryst (i.e., phases

plagioclase, pyroxenes, Fe-Ti oxides, olivine). In fact, FeO<sub>t</sub>, TiO<sub>2</sub>, CaO, and Al<sub>2</sub>O<sub>3</sub> are positively correlated with MgO, while SiO<sub>2</sub> and Na<sub>2</sub>O are negatively correlated with MgO. The extent of Al<sub>2</sub>O<sub>3</sub> and CaO scattering, however, implies at least different mineral phase proportions and/or compositions during fractional crystallization.

Trace elements depict broad correlations with silica, despite considerable scattering. In particular, Rb, Ba, Zr, Y and Nb (TABLE 6, Fig. 6) exhibit an incompatible behaviour, as also evidenced by their mutual positive correlations (not shown). Conversely, V and Sr behave as compatible elements and reflect the progressive removal of Fe-Ti oxides and plagioclase. Cr and Ni display low concentrations (<100 and <60 ppm, respectively; TABLE 6), a typical feature of low-MgO, high-Al<sub>2</sub>O<sub>3</sub> arc magmas worldwide, which is related to the prolonged removal of mafic phases during magma differentiation (e.g., Pichavant and Macdonald, 2007 and references therein).

Spider diagrams of trace element concentrations normalized to N-MORB (Sun and Mc Donough, 1989; Fig. 7a), show depletion of HFSE (High Field Strength Elements; with pronounced troughs at Nb and Ti) with respect to LILE (Large Ion Lithophile Elements) and LREE (Light Rare Earth Elements), as it is typical of arc magmas. The whole set of samples shares very similar patterns, with a general enrichment toward the most differentiated rocks and only slight changes in the incompatible/compatible element ratios. The chondrite-normalized REE patterns are also similar for rocks with different degree of evolution (Fig. 7b) and show a slight increase of  $(La/Yb)_N$  (from 3.6 to 5.9; subscript N indicates normalized abundances), for near constant  $(Sm/Yb)_N$ , in the most differentiated rocks. The latter also show a weak Eu trough (Eu/Eu\*=0.83-0.88, where Eu\* is (Sm+Gd)/2, which instead is lacking in less differentiated rock types (Eu/Eu\*=0.93-0.97).

				Olivine			
Unit	SAV12	SAV12	SA7 LT	SA35 LT	SA35 LT	SA40 LT	SA40 LT
	c	mph	c	c	r	c	r
SiO <sub>2</sub>	35.73	38.70	35.53	37.56	38.58	36.7	38.13
FeO	37.82	32.27	37.86	34.76	36.1	37.7	39.047
MnO	0.64	0.68	0.75	0.52	0.24	0.64	0.50
MgO	26.37	25.94	24.99	25.91	24.18	24.94	22.06
CaO	0.64	0.71	0.43	0.47	0.70	0.29	0.41
Total	100.56	97.58	99.13	98.75	99.10	99.98	99.74
Fo (mol%)	55.42	58.90	54.06	57.06	54.43	54.12	50.18
		Amphibole				Fe-Ti oxides	5
	SA34	SA25	SA32		SAV8	SAV8	SAV11
Unit	LI	F	LI		LI	LI	LT
SiO <sub>2</sub>	47.53	47.71	47.41		-	0.21	-
TiO <sub>2</sub>	1.44	1.34	1.45		1.70	45.84	3.12
$Al_2O_3$	8.68	8.33	8.18		1.43	0.12	1.34
FeO	14.58	13.61	13.79		89.99	50.32	88.39
MnO	0.07	0.12	0.12		0.35	2.13	0.37
MgO	13.75	14.41	14.37		-	0.1	0.33
CaO	10.34	10.54	10.97				
Na <sub>2</sub> O	1.36	1.61	1.37				
K <sub>2</sub> Ô	0.21	0.31	0.36				
Total	97.96	97.98	98.02		93.47	98.72	93.55
mg#	0.63	0.65	0.65				

 TABLE 5

 Representative electron microprobe analyses of olivine, amphibole and Fe-Ti oxides from the Sant'Antioco study rock-types.

c: phenocryst core ; r: phenocryst rim; mph: microphenocryst; mg# = MgO/(MgO+FeO) molar, with all iron as  $Fe^{2+}$ .

Overall, trace element patterns indicate that magmatic evolution was mostly controlled by fractional crystallization, consistent with evidence from major element variations. The wide data point scatter, however, implies that fractional crystallization cannot fully account for the observed chemical variations in the Sant'Antioco volcanics and that an additional process, such as crustal contamination, was likely involved in magma evolution (also see below). This seems to be confirmed by the <sup>87</sup>Sr/<sup>86</sup>Sr initial (i.e. at 15 Ma) isotopic ratios for Sant'Antioco whole rocks, which range from 0.70698 to 0.70858±1 (TABLE 6). Crustal assimilation is also suggested by the positive correlations of the initial Sr isotope ratios with



Fig. 4 - Classification diagrams of  $K_2O vs$  SiO<sub>2</sub> (after Peccerillo and Taylor, 1976) (a), MgO vs Al<sub>2</sub>O<sub>3</sub> (after Kersting and Arculus, 1994) (b), FeO\*/MgO vs SiO<sub>2</sub> (after Miyashiro, 1974) (c), for the investigated Sant'Antioco volcanics. Inset in (c) shows the plot of FeO\* vs FeO\*/MgO (\* = all iron as FeO). The three samples representative of the least evolved compositions (i.e., SA7, SA35 and SA40) are marked. Solid lines enclose the compositional field of Cenozoic orogenic volcanic rocks of Sardinia (data from literature cited in the text). HMB: high-magnesia basalt; HAB: high-alumina basalt;  $\alpha$ : medium-K basalt;  $\beta$ : medium-K basaltic andesite;  $\gamma$ : medium-K andesite;  $\delta$ : medium-K dacite.

 $SiO_2$  and with some incompatible trace elements sensitive to crustal contamination (i.e., Rb, Ba, Zr), along with the negative correlation with Sr abundance, sensitive to crustal contamination + plagioclase fractionation (Fig. 8).

# Estimates of magma temperatures and $H_2O$ concentrations

Geothermometric determinations based on magnetite/Ti-magnetite and ilmenite pairs (Buddington and Lindsley, 1964) are prevented for the Sant'Antioco suite, due to the rare

occurrence of coexisting, unexsolved crystals of the two opaque phases in the analyzed samples. Therefore. in order to constrain the crystallization temperatures of magmas, the twopyroxenes equilibrium geothermometer of Lindsley (1983) was applied. This approach may provide a general indication of temperature even in the lack of coexisting low-Ca pyroxene, by individual augite plotting phenocryst compositions. The projection onto Lindsley's isotherms for P = 500 MPa (preferred to the 1 GPa and 1 atm schemes to constrain the maximum crystallization depth at Sant'Antioco,

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Representative analyses of major and trace elements from Sant'Antioco study rock-types. <sup>87</sup>5/87 isotope ratios of selected samples are also reported.

		AND	57.22 0.70 19.71 3.39	3.36 0.13	1.78 8.57 2.78	0.95	99.96 6.41 33.1	5202 5333364 533364 53364 5207
(J	AV16**	AND	56.00 : 0.91 18.48 3.63	4.53 0.16	2.18 8.44 8.18	0.17	99.94 7.80 37.2	154 19 9 9 9 379 458 458 458 458
S UNIT	SAV17**S	HABA	55.55 0.90 18.96 3.82	4.27 0.16	2.58 8.69 2.48	0.17	99.96 7.71 37.4	142 15 15 15 7 7 7 7 7 7 3 3 9 3 9 3 9 3 9 3 9
ARBU	SAV19**S	HABA	54.51 0.96 18.70 3.69	0.17	3.07 8.63 2.30	0.20	99.96 8.31 39.7	133 12 133 374 28 357 357 357 357 357
A - M.te	SAV14*S	HABA	53.91 0.99 19.47 4.56	4.06 0.17	3.19 9.44 2.53	0.17	99.97 8.16 8.11	148 148 20 25 25 2.6 2.6 2.6 2.6 2.6 2.6 2.6
ISPISI/	SAV15*	HABA	53.18 1.05 19.17 4.42	4.59 0.19	3.70 9.24 2.49	0.16	99.96 8.57 8.57	$\begin{array}{c} 180\\ 180\\ 180\\ 180\\ 180\\ 112\\ 180\\ 112\\ 180\\ 112\\ 180\\ 112\\ 180\\ 112\\ 180\\ 112\\ 180\\ 122\\ 1122\\ 1122\\ 1226\\$
Ч	SA27*	HABA	52.68 1.07 18.82 3.43	5.53 0.16	9.37 2.30	0.16	99.94 8.62 8.7.7	162 102
	SA32	DAC	5.04 0.60 5.33 4.27	0.13	1.74 5.56 2.55	0.19	9.98 5.95 4.3	83 85 55 55 55 55 55 7 7 7 7 7 7 2 55 55 3 55 55 55 55 55 55 55 55 55 55 5
	A33 S	ND I		12	66 66	19	100 100 100 100 100 100 100 100 100 100	055 5 4 5 6 6 6 7 7 3 7 5 7 5 7 5 7 5 7 5 7 5 7 5 7 5 7
	.34 S/	ND AI	0.33 61 0.71 0 0.71 17 17 17 17 17	13 0		22 0		500005665054 5000000 5000000 500000 500000 500000 500000 500000 500000 500000 500000 500000 500000 500000 50000 50000 50000 50000 50000 5000 5000 5000 50000 5000 5000
	NS S/	ND AI	0.51 60 0.75 0 7.36 17 4.04 4	0.15	05.30 74 50 77 6	0.21	2007 99 2019 99 2019 10 2019 1	2 4 8 2 2 2 2 3 8 2 8 2 7 5 7 5 7 5 7 5 7 5 7 5 7 5 7 5 7 5 7
	V5 S/	ND A	200 59 200 50 200 50 50 50 50 50 50 50 50 50 50 50 50 50 5	16 11 2	202	523		88 88 88 88 88 88 88 88 88 88 88 88 88
	17 SA	N A	00 58 29 17 60 3	19 0	01 2 4 5 01 3 1 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7	28 0	.07 99 .97 99 .85 6 .7 36	
	18 SA	D AN	73 58 69 1 62 17 95 4	65 112 0 21 0 21 0 21 0 21 0 21 20 20 20 20 20 20 20 20 20 20 20 20 20	49 96 3 7 2	89 1 25 0 1	97 99 97 99 9 35	885543885999554 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 88554388599955 8855438859955 8855438859955 8855438859955 8855438859955 8855438855 8855438855 8855438855 8855438855 8855438855 8855438855 8855438855 8855438855 885545 885545 885545 885545 885545 885545 885555 885545 885555 885555 885555 885555 885555 885555 885555 8855555 885555 885555 885555 885555 885555 8855555 8855555 8855555 8855555 8855555 8855555 8855555 8855555 88555555
	I SAV	ANI (	0. 19. 19. 19. 19. 19. 19. 19. 19. 19. 19	04	×222	6000	07 99. 17 6. 132.	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
	SA3	AND	2 57.0 3 0.8 18.4 0 5.0	0.00			5.99.9 9.99.9 39.0	111 111 188 488 4357 4357 4357 4357
([]	SA10	AND	56.8 0.9 18.0 2.90	5.3 0.1	5 7 8	0.5	99.99 7.9	$\begin{array}{c} 191\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 8\\ 7\\ 7\\ 7\\ 7\\ 7\\ 7\\ 7\\ 7\\ 7\\ 7\\ 7\\ 7\\ 7\\$
VAS (I	SA15	AND	55.90 0.90 18.51 3.92	0.16	3.18 8.57 2.60	0.23	99.98 99.98 7.76 42.4	186 122 122 139 139 139 139 139 139 139 139 139 139
TY LA	SAV13	AND	55.28 0.90 19.09 3.36	4.55	2.12 8.67 2.62	0.21	99.98 99.98 7.57 33.3	191 6 7 357 357 7 405 4.5 4.5
EAR	SA12	AND	55.19 0.96 19.82 4.08	3.19	9.02 2.62	0.18	99.97 6.86 36.3	180 9 13 13 13 13 13 14 8 8 448 448 448
	SA11	HABA	54.89 1.03 18.94 3.61	4.78 0.16	3.23 8.63 2.65	0.20	99.95 99.95 8.03 41.7	$\begin{array}{c} 222\\ 10\\ 202\\ 202\\ 202\\ 202\\ 202\\ 202\\ $
	LN	ABA	4.60 9.43 9.43	3.45	2.75 2.75	0.19	0.97 9.97 7.95 6.9	212 19 19 33 33 33 33 33 33 33 35 35 35 37070 07070
	S	Η	ν <del>-</del> -				6 6	

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		SA25 AND	60.60	0.85	16.79	5 13	0.17	1.57	5.78	3.55	1.94	0.28	1.20 99.95	6.92	28.8	77	t 2	14	5	67	314	35	10	482	6.6	29.9 62 4	7.12	32.3	6.69	1.78	0.00	5 94	1.19	3.47	0.52	3.36	0.52	4.03 0.01	1.13	4
	S (F)	SAV8b AND	58.93	0.68	17.99	3.64 2.40	0.14	2.00	7.13	2.80	1.57	0.19	1.41 99.96	6.78	34.5	511	111	15	12	57	280	24	171	539	5.6	22.9	5.12	22.4	4.99	1.68	4.86	4 71	1.03	2.99	0.44	3.01	0.45	4.31 0.87	0.98	$0.708^{4}$
	DYKE	SA21 AND	55.26	1.04	18.26	4 49 4 49	0.15	2.80	8.97	2.33	0.98	0.20	06.96	8.29	37.6	175	19	18	7	35	378	27	100	520	4.4															
		SA29 HABA	52.58	1.12	18.19	4 07	0.18	4.19	9.45	2.44	0.69	0.22	1.40 99.96	8.91	45.6	120	107	22	19	23	337	24	901	378	2.3															
		SA9 HABA	54.38	1.03	19.66 2 15	0.4.0 4.40	0.16	3.03	8.89	2.53	1.00	0.19	99.96	7.50	41.9	200	15	19	7	34	302 2	26 176	07 X	354	4.1															
		SA38 HABA	54.14	0.90	20.80	06.7 15 4	0 14	2.17	9.68	2.65	0.93	0.23	0.95 09.95	7.17	35.0	105	107 96	15	15	25	368	78	101	310	3.1	15.9 22.1	4.13	20.2	4.74	1.42	4.58	4 74	0.95	2.86	0.42	2.74	0.42	71.7	0.58	
6 6 ed		SA1 HABA	53.70	1.00	26./1	3.30	0.18	4.00	8.79	2.46	1.17	0.23	c0.1 99.63	8.49	45.6	010	249 31	27	15	33	340 22	52	101	301	2.8															
TABLE	~	SA3 HABA	53.23	1.00	18.36	00.0 4 98	0.18	3.85	9.16	2.59	1.19	0.22	99.95 95	8.40	45.0	370	707 19	23	14	28	355	22	201	343	2.7															
0	AS (LT	SA36 HABA	53.16	1.00	19.00 2 00	5.00 5.18	0.18	3.92	9.15	2.45	0.97	0.20	0.80 99.95	8.67	44.6	750	164	18	34	25	315 2	50 07	6 4	291	2.6	13.2	3.35	16.7	3.83	1.24	4.20	0.70	0.89	2.70	0.39	2.52	0.39	14.7	0.54	98
	AL LAV	SAV11 HABA	52.65	1.03	20.32	4.42	0.16	2.68	10.11	2.46	1.03	0.21	0.81 99.95	8.05	37.2	"	91	25	20	26	362	73	64	397	2.9															0.706
	FIN/	SA35 HABA	51.68	1.03	20.68	4.07 ۲07	0.16	3.09	10.46	2.36	0.57	0.21	96.96	8.10	40.5	151	707	17	35	22	398 22	52	200	345	3.2	13.4	20.2 3.46	17.3	3.96	1.29	4.25	4 40	0.89	2.68	0.42	2.50	0.40	0.44	0.37	
		SA7 HABA	51.13	1.04	18./8	00.4 707	0.18	4.61	9.87	2.07	0.69	0.16	1.85 99.95	9.12	47.4	000	100	28	41	21	336	57 0	₽ 4	380	2.4	13.9	3.49	17.0	3.88	1.24	4.19	4 15	0.83	2.55	0.35	2.28	0.35	17.7	0.49	
		SA40 HABA	51.12	1.07	19.64 5 00	4 48	0.18	3.56	9.90	2.49	0.96	0.18	16.1 99.97	9.05	41.2		767	27	28	27	352	528	707	322	2.8	13.4 0.70	3.29	16.6	3.87	1.24	4.08	4 11	0.83	2.47	0.37	2.35	0.35	C7.7	0.52	16
		SAV12 e HAB	50.86	1.18	20.23	4.4 	0.16	2.78	10.65	2.37	0.76	0.17	1.04 99.94	9.29	34.8	020	617	21	16	23	338	23	ç (,	402	1.9	13.6	3.24	16.8	3.67	1.24	3.99	4.08	0.84	2.39	0.35	2.25	0.36	c/ .1	0.41	r) <sub>i</sub> 0.707
		Sample ªRock-typ	$SiO_2$	TiO <sub>2</sub>	${\rm AI_2O_3}$	FeO3	MnO	MgO	CaO	$Na_2O$	$\rm K_2O$	$P_2O_5$	г.О.Т	$FeO_{nul}$	mg#	11	۔ ح ح	රී	Ņ	Rb	Sr	Y Y	z fz	Ba	Th	La	Pr Pr	PN	Sm	Eu	5f		Ho H	Er	Tm	Yb	Lu	ΗĽ	U U	( <sup>87</sup> Sr/ <sup>86</sup> S

HAB: high-alumina basalt; HABA: highlumina basaltic andesite; AND: andesite; AAC: dacite. lava injection in polygenic breccia; \* lava clasts in polygenic breccia; ng# = MgO/(MgO+FeO\*) molar, with all ron as FeO.



Fig. 5 - Variation diagrams of selected major oxides vs MgO for the investigated Sant'Antioco volcanic suite. The three samples representative of the least evolved compositions (i.e., SA7, SA35 and SA40) are marked.



Fig. 6 - Variation diagrams of selected trace elements vs SiO<sub>2</sub> for the investigated Sant'Antioco volcanic suite. The three samples representative of the least evolved compositions (i.e., SA7, SA35 and SA40) are marked.

according to the observed phase relationships and MELTS modeling; Ghiorso *et al.*, 1994; see below in this Section) indicates that most temperatures fall in the range of 1000-1150°C. One of the least evolved basaltic-andesites, SA7 (LT), shows the highest equilibration temperature at  $\sim$ 1200°C (Fig. 9).

In addition, crystallization temperatures were



Fig. 7 - Normal MORB-normalized diagram of trace element concentrations (a) and chondrite-normalized REE patterns (b) for representative Sant'Antioco volcanics (HABA: high-alumina basaltic andesite; AND: andesite).

estimated by applying the plagioclase-liquid geothermometer (Putirka, 2008) to the plagioclase cores-whole rock pairs. For these calculations, basaltic andesite and andesite samples characterized by the lowest contents of phenocrysts (either isolated or in glomeri) were selected as the best representative of liquid compositions. The calculated temperature interval of 1100-1165°C broadly matches that obtained through Lindsley's (1983) method.

The amount of dissolved  $H_2O$  in the system can be also estimated based on the plagioclasemelt equilibrium equation (Putirka, 2008). Calculations yield  $H_2O$  concentrations in the 2.3-3.9 wt% range, where the highest value refers to the highest- $Al_2O_3$  lava sample (SA35). A similar range of  $H_2O$  contents (2.0-4.1 wt%) is derived through the Pichavant and Macdonald (2007) equation (an empirical geohygrometer based on melt compositions for arc basalt melts), using T values obtained from pyroxene geothermometry. Notably, among the least evolved lavas (mg#  $\geq$ 0.4), SA35 records both the highest  $H_2O$  content (~4 wt%) and the lowest crystallization temperature (TABLE 7).

The obtained  $H_2O$  contents are consistent with the observed plagioclase compositional variations, the An/Ab ratio being sensitive to  $X_{H_{2}O}$  in the melt. From experimental evidence (Sisson and Grove, 1993a, b), the  $K_D^{Ca-Na}$  (=  $X_{Ca}^{Pl}$ \*  $X_{Na}^{liq} / X_{Na}^{Pl}$  \*  $X_{Ca}^{liq}$ ) for plagioclase-melt equilibria of high-alumina basalts and basaltic andesites is strongly dependent on H<sub>2</sub>O contents and relatively independent on pressure, as it increases from 1.1 (anhydrous conditions) up to 5.5 (H<sub>2</sub>O-saturated magmas).

On these grounds, the plagioclase-liquid Ca/Na (molar) diagram provides constraints on the pre-eruptive  $X_{H_{2}O}$  conditions of the studied products (Fig. 10). Concerning high-alumina basaltic andesites, K<sub>D</sub>Ca-Na values for microphenocrysts and phenocryst rims vs bulk groundmass compositions range from ca. 1.3 to 2.3, most of the data points plotting between the experimental K<sub>D</sub><sup>Ca-Na</sup> equilibrium lines for anhydrous (K<sub>D</sub><sup>Ca-Na=1.1</sup>) and 2 wt% H<sub>2</sub>Ohydrated melts (K<sub>D</sub><sup>Ca-Na=1.7</sup>; Fig. 10). Phenocryst cores vs whole rock equilibria yield higher  $K_{D}^{Ca-Na}$  values (from ca. 2.6 to 4.5), corresponding to the interval between 2 wt%  $H_2O$  ( $K_D^{Ca-Na=1.7}$ ) and the  $H_2O$ -saturated conditions (6 wt%  $H_2O$ ;  $K_D^{Ca-Na=5.52}$ ). Phenocryst cores vs whole rock data for selected andesites (Fig. 10) reveal lower H<sub>2</sub>O concentrations than in basaltic andesites



Fig. 8 - Variation diagrams of <sup>87</sup>Sr/<sup>86</sup>Sr vs SiO<sub>2</sub> and selected trace elements for representative Sant'Antioco mafic volcanics. Data for volcanic rocks from Montresta (HMB: high-MgO basalt, HAB: high-Al<sub>2</sub>O<sub>3</sub> basalt, BA: basaltic andesite and AND: andesite) and Sarroch (basaltic andesite and andesite) districts (after Morra *et al.*, 1997 and Conte, 1997) are also reported for comparison.



Fig. 9 - Representative analyses of pyroxenes from the Sant'Antioco study suite, projected onto the Di-En-Fs-Hd quadrilater (following Lindsley, 1983). Isotherms are shown for P = 500 MPa, chosen, in preference to 1 GPa or 1 atm, as the most appropriate pressure conditions for magma crystallization in our case study (see text).

TABLE 7 Calculated water contents ( $H_2O$  wt%) for selected Sant'Antioco samples, following the methods of Pichavant and McDonald (2007) and Putirka (2008).

Rock samples	Pichavant and Macdonald (2007)	Putirka (2008)
SAV12	3.8	-
SA7	2.6	2.4
SA35	4.1	3.9
SA40	2.5	2.7
SAV11	2.9	2.5
SA15	3.1	-
SAV15b	3.4	2.6
SAV13	3.3	-
SA10	2.4	-
SAV10	3.1	-
SAV5	2.0	2.3
SAV8	3.5	-

(comparable to the values obtained from crystal rims in these latter).

Furthermore, the relationships between the calculated H<sub>2</sub>O contents in the melts and plagioclase compositions for some of the least evolved lavas (i.e., SA7, SA40 and SA35) were checked following silicate melt modeling (Ghiorso et al., 1994) for different amounts of dissolved water in the system (1-6 wt%), within the load pressure interval of 100-400 MPa. Model results broadly confirm the above estimated water contents, the latter being the most critical parameter controlling K<sub>D</sub> values (rather than mere load pressure). For example, a run at P=100 MPa, T=1110°C and 2 wt% H<sub>2</sub>O fits the calculated K<sub>D</sub><sup>Ca-Na</sup> for phenocryst cores vs whole rock for sample SA7, while a run at P=200 MPa, T=1040°C and 5 wt% H<sub>2</sub>O yields a  $K_D^{Ca-Na}$  matching that obtained for sample SA35. Finally, from pilot experiments performed on the



Fig. 10 - Comparison of [Ca]/[Na] (molar) in plagioclase and coexisting liquid for selected samples from Sant'Antioco volcanics (HABA: high-alumina basaltic andesite; AND: andesite). The three samples representative of the least evolved compositions (i.e., SA7, SA35 and SA40) are marked. Each bar represents the range of values for a single rock-sample. Crystal-liquid equilibria are referred to either phenocryst cores *vs* whole rock compositions or phenocryst rims and microphenocrysts *vs* bulk groundmass compositions. Straight lines represent experimental equilibrium  $K_D^{Ca-Na}$  values for anhydrous ( $K_D^{Ca-Na}$ =1.1) to H<sub>2</sub>O-saturated ( $K_D^{Ca-Na}$ =5.5) melts after Sisson and Grove (1993a).

Sant'Antioco least evolved rocks, it has been inferred that SA35 experienced crystallization under significantly higher  $P_{H_{20}}$  with respect to SA40 and SA7 (Conte *et al.*, 2000).

# PETROGENETIC INFERENCES

The paucity of primitive magma compositions in subduction-related volcanism worldwide precludes the assessment of the origin of primary magmas, and makes it difficult to distinguish the effects of subcrustal vs crustal processes. Similarly, as concerns the rather differentiated lavas from Sant'Antioco, only low-pressure processes involved in their petrogenesis can be addressed here. Indications on source-related processes can be only tentatively drawn by comparison with less-evolved Cenozoic calcalkaline rocks of Sardinia.

By considering the new set of mineropetrographic and geochemical data, in the light of the huge amount of literature studies on highalumina arc magmatism, it appears that the wide spectrum of low-MgO (and low Mg#, Ni, Cr), high-alumina rock-types from Sant'Antioco likely reflects extensive fractional crystallization of a parental basaltic liquid, where changes in the fractionating mineral assemblages occurred in response to variable PH2O conditions (Sisson and Grove, 1993a, b; Pichavant and Macdonald, 2007 and references therein; Ulmer, 2007 and references therein), coupled with crustal interaction at some extent (as suggested by trace element and Sr isotope data; Geochemistry section). On the other hand, the lack of typical disequilibrium textures (i.e., rim resorption, mantling by other phases, reverse or oscillatory zoning) in pyroxene crystals does not support additional low-pressure processes, such as mixing between compositionally different magmas. In this regard, plagioclase dusty textures and zoning patterns are not considered as conclusive evidence of magma mixing and are, instead, thought to be essentially controlled by H<sub>2</sub>O conditions (also cfr. Lonis et al., 1997). Even if, in some cases, the observed plagioclase disequilibrium textures and the occurrence of polycrystalline aggregates in LI andesites may point out some crystal recycling in the magma reservoir, we exclude that this process may have significantly affected the chemical evolution of the Sant'Antioco magmas. In this regard, the observed wide scattering in Al<sub>2</sub>O<sub>3</sub> and CaO vs MgO contents (Fig. 5) would imply unlikely high amounts (some tens vol%) of plagioclase+pyroxenes-bearing glomeri (also cfr. Brotzu et al., 1997a), which do not actually occur in the study volcanics. Moreover, trace element patterns (e.g., the absence of Sr, Eu and Cr positive anomalies) provide no evidence of significant cumulitic processes.

In order to constrain in more detail the petrogenetic processes involved in the studied rock suite, we focus now on the least evolved compositions (i.e., SA7, SA40 and SA35, highalumina basaltic andesites). which are characterized by different MgO, Al<sub>2</sub>O<sub>3</sub>, K<sub>2</sub>O and FeO contents, in spite of similar SiO<sub>2</sub> contents (around 52 wt%). In Fig. 5, the Al<sub>2</sub>O<sub>3</sub>-richest sample, SA35, also displays slightly higher CaO and lower FeO contents with respect to the MgO-richest sample, SA7, while sample SA40 shows intermediate contents. Trace element contents are quite similar, except for higher Sr in SA35 and slightly higher Ni and Cr in SA7 (TABLE 6).

In addition, the three samples are characterized by distinctive phenocryst assemblages, although all plagioclase-dominated (TABLE 2): the Al<sub>2</sub>O<sub>3</sub>richest, phenocryst-poorest, sample SA35 contains the lowest plagioclase amount, and olivine plus rare Cpx as mafic phases; SA40, with intermediate Al<sub>2</sub>O<sub>3</sub> content, shows higher abundances of plagioclase and mafic phases, with olivine again prevailing on Cpx; in SA7, the Al<sub>2</sub>O<sub>3</sub>-poorest and MgO-richest sample, Cpx prevails on olivine, and minor amounts of Opx are also observed. The three samples also differ in mineral chemistry, being characterized by distinct plagioclase compositions (i.e.  $An_{75-85}$ ,  $An_{78-90}$  and  $An_{85-93}$  for SA7, SA40 and SA35, respectively; Fig. 2) and different Al<sub>2</sub>O<sub>3</sub> contents in Cpx (from ~2wt% in SA7 up to ~6wt% in SA35), which parallel whole-rock Al<sub>2</sub>O<sub>3</sub>.

These differences likely indicate that corresponding magmas underwent different preeruptive conditions storage of and crystallization, consistent with the above estimated values of temperature, pressure of crystallization and dissolved water content for the three cases. In particular, plagioclase-host liquid equilibria (Fig. 10) provide quantitative information on H<sub>2</sub>O contents in the three magmas at different evolutionary stages (the liquid compositions being roughly represented by either the whole-rock or groundmass bulk compositions with ongoing magma crystallization).

Water content data broadly fit the abovereported mineralogical features of samples. In fact, it has been proven that high dissolved  $H_2O$ concentrations in high-alumina melts, not only favour the early formation of spinel and the crystallization of An-richer plagioclase, but also reduce the total proportion of plagioclase in the crystallizing assemblage, expand the stability field of olivine at the expense of other mafic phases such as Cpx and Opx, and decrease the En-Fs contents of Cpx in favour of CaTs (Kushiro, 1969; Sisson and Grove, 1993a; Pichavant and Macdonald, 2007).

These considerations lead us to conclude that changes in phenocryst proportions and compositions (and consequent changes in bulk rock chemistry) in the least evolved basaltic andesites were determined by different  $H_2O$ concentrations in the melts and by the physical conditions controlling water exsolution from the system during magma ascent. In summary: the  $Al_2O_3$ -richest, MgO-poorest sample (SA35) displays, consistent with the highest  $X_{H_2O}$  (up to ca. 6 wt%, Fig.10), the lowest modal amount of plagioclase, accompanied by the Ca-richest composition, a paucity of Cpx, with the Ca-Tsrichest augite composition, and the presence of olivine. Conversely, the MgO-richest, Al<sub>2</sub>O<sub>3</sub>poorest composition of SA7, may reflect a crystallization path under H<sub>2</sub>O-poorer conditions (up to ca. 3 wt%), which favoured the crystallization of higher amounts of Ca-poorer plagioclase and Al<sub>2</sub>O<sub>3</sub>-poorer Cpx, as well as the occurrence of Opx at the expense of olivine. SA40, which shows intermediate Al<sub>2</sub>O<sub>3</sub> and MgO contents, consistently accounts for intermediate H<sub>2</sub>O concentration in the magma (up to ca. 4 wt%). In this picture, the Ab-richer plagioclase rims and microphenocrysts in the three samples may testify late H<sub>2</sub>O exsolution.

Moreover, modes of water retention/exsolution in the system influence the amphibole stability field. Notably, amphibole only occurs in two LI samples, which records prolonged  $H_2O$  retention in the corresponding magmas; late-stage processes of decreasing pressure and  $H_2O$ exsolution, causing amphibole break-down, are testified by the presence of opacitic rims.

The effect of  $X_{\rm H_{2O}}$  on phase equilibria thus may represent the key factor in determining the observed geochemical differences among the least evolved Sant'Antioco rock-types, as a result of different evolution paths from a possible common parental melt.

# Identification of the parental magma

As it is typical of arc-related magmatism worldwide (e.g., Nye and Reid, 1986; Brophy, 1989; Eggins, 1993; Sisson and Grove, 1993b; Rohrbach *et al.*, 2005; Pichavant *et al.*, 2002; Pichavant and Macdonald, 2007), in Cenozoic orogenic magmatism of Sardinia, high-magnesia basalts (HMBs, MgO >8 wt%), possibly representative of mantle-derived magmas, are relatively uncommon and limited to a few areas (Mattioli *et al.*, 2000 and references therein). In

particular, HMBs from the Montresta district, NW Sardinia (18.5 Ma, Montigny et al., 1981; Morra et al., 1997; Fig. 1) could match in composition the possible precursor for Sant'Antioco LT basaltic andesites. In fact, the high-Al<sub>2</sub>O<sub>3</sub> basaltic andesites derived from Montresta **HMBs** through fractional crystallization±assimilation processes (Morra et al., 1997) show major and trace element compositions that closely resemble those from Sant'Antioco (Fig. 11, 12), except for significantly lower Ba contents (e.g., 140-250 vs 300-400 ppm, for Montresta and Sant'Antioco HABA, respectively; Fig. 8).

If a typical Montresta HMB (e.g., sample KB13, MgO = 9.08 wt%, Morra *et al.*, 1997) is chosen as representative of the Sant'Antioco parental magma composition, major element mass balance calculations (following Stormer and Nicholls, 1978) broadly indicate that samples SA7, SA40 and SA35 may be consistently derived by fractional crystallization under different  $P_{H_{2O}}$  conditions (TABLE 8). Notably, the results yield similar fractions of residual liquid for the three cases (~45-50 wt%), but different proportions of the main crystal phases (especially plagioclase *vs* mafic phases), i.e.: Plg/(Ol+Cpx)=0.52, 0.36 and 0.24 for SA7, SA40 and SA35, respectively.

The similar fractions of residual liquid may account for the similar contents in the majority of incompatible trace elements in the three cases, whereas the observed differences in the Sr contents can be explained by the different amounts of plagioclase fractionation (Fig. 6; 7a, b).

It is also worth noting that relatively high  $X_{H_{2}O}$  conditions *ab initio*, which determine the fractionation of high proportions of mafic phases (with high MgO/FeO\* ratios), with respect to plagioclase, may even result in the observed ambiguous tholeiitic character of the Sant'Antioco mafic volcanics (Fig. 4). In our interpretation, this inference could also apply to the "tholeiitic" lavas in the Montresta district

#### (Morra et al., 1997).

#### Magma evolution

From the whole compositional spectrum of the studied Sant'Antioco volcanics it appears that the differences existing in the least evolved compositions also reflect in the differentiation products (Fig. 5, 11). Major-element massbalance calculations (TABLE 9) show that the latter can be derived through prolonged fractional crystallization of Plg + Cpx + Fe-Ti oxides  $\pm Ol \pm Opx$  mineral assemblages. In more detail, if SA35 and SA7 lava-types, which display the most significant differences in terms of oxide concentrations and H<sub>2</sub>O contents, are considered as starting melts, the results of calculations evidence that olivine continues to play a role in the early stages of the H<sub>2</sub>O-rich differentiation trend starting from SA35 (Fig. 11), while Opx may replace olivine in the fractionating crystal assemblage along the H<sub>2</sub>Opoor differentiation trend starting from SA7 (Fig. 11, TABLE 9). These results confirm, also in agreement with experimental evidence (Pichavant and Macdonald, 2007 and references therein), that in mafic to intermediate calcalkaline magmas Opx is stable only under low H<sub>2</sub>O concentrations, while in more hydrous conditions Opx crystallization is delayed and takes place extensively only in SiO<sub>2</sub>-richer magmas at lower temperatures. Indeed, in the Sant'Antioco volcanics, like in other typical arcrock suites (e.g., Macdonald et al., 2000), Opx is always present in the most differentiated rocks  $(SiO_2 > 55 \text{ wt\%}; \text{ TABLE } 9)$ . In more hydrous melts, Opx crystallization is joined by amphibole during the late differentiation stages.

However, simple fractional crystallization cannot fully explain the whole compositional spectrum of the study rock suite. In this regard, the above reported <sup>87</sup>Sr/<sup>86</sup>Sr data (0.70698 to 0.70858±1; TABLE 2; Fig. 8) are in the 0.70398-0.71130 range pertaining to the Cenozoic

Modeled fractionation steps (following Stormer and Nicholls, 1978) for the least evolved basaltic andesites from Sant'Antioco, starting from a typical high-magnesia basalt of Montresta, inferred as possible parental magma. TABLE 8

		from KB	13 to SA35	from KB	13 to SA40	from KE	313 to SA7
	KB13*	SA35	residuals	SA40	residuals	SA7	residuals
SiO <sub>2</sub>	48.88	52.54	0.02	52.06	0.00	52.34	0.04
TiO <sub>2</sub>	0.75	1.04	0.00	1.09	0.01	1.07	0.01
$Al_2O_3$	15.55	21.02	0.00	20.00	0.00	19.22	0.02
FeO*	10.16	8.27	0.02	9.27	0.04	9.38	0.05
MnO	0.12	0.16	0.00	0.18	0.02	0.18	0.01
MgO	9.08	3.14	0.00	3.63	0.02	4.72	0.00
CaO	12.90	10.63	0.01	10.08	0.01	10.11	0.01
$Na_2O$	1.81	2.40	0.14	2.53	0.05	2.12	0.26
$K_2O$	0.52	0.58	0.05	0.98	0.00	0.70	0.04
$P_2O_5$	0.23	0.21	0.02	0.18	0.02	0.16	0.03
	100.00	100.00		100.00		100.00	
$\sum r^2$			0.25		0.16		0.48
			Remo	oved crystals (v	vt%)		
01		11.89		10.46		8.17	
Cpx		28.41		28.18		27.51	
Plg		9.57		13.74		18.66	
Fe-Ti oxides		0.10				0.75	
wt% of residual liquid		50.03		47.62		44.91	
Cpx/Ol		2.39		2.70		3.37	
Cpx+OI		40.30		38.63		35.68	
Plg/(Cpx+Ol)		0.24		0.36		0.52	

SA7         SA15b         residuals           SiO2         52.34         56.68         0.00           TiO2         1.07         0.95         0.12           A1 <sub>2</sub> O3         19.22         18.43         0.00           FeO*         9.38         8.11         0.00           MnO         0.18         0.15         0.00           MgO         4.72         3.38         0.00           CaO         10.11         8.32         0.00	SA32b			Εī	om SA35 to	SA16 to SAV	18	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$		residuals		SA35	SA16	residuals	SAV18	residuals
TIO2         1.07         0.95         0.12           Al <sub>2</sub> O3         19.22         18.43         0.00           FeO*         9.38         8.11         0.00           MnO         0.18         0.15         0.00           MgO         4.72         3.38         0.00           CaO         10.11         8.32         0.00	60.88	0.01	$SiO_2$	52.54	54.50	0.01	58.49	0.04
Al <sub>2</sub> O <sub>3</sub> 19.22 18.43 0.00 FeO* 9.38 8.11 0.00 MnO 0.18 0.15 0.00 MgO 4.72 3.38 0.00 CaO 10.11 8.32 0.00	0.73	0.10	$TiO_2$	1.04	1.05	0.02	0.70	0.17
FeO*         9.38         8.11         0.00           MnO         0.18         0.15         0.00           MgO         4.72         3.38         0.00           CaO         10.11         8.32         0.00	17.57	0.00	$Al_2O_3$	21.02	20.20	0.00	19.88	0.02
MnO         0.18         0.15         0.00           MgO         4.72         3.38         0.00           CaO         10.11         8.32         0.00	6.81	0.01	FeO*	8.27	7.96	0.00	6.32	0.04
MgO 4.72 3.38 0.00 CaO 10.11 8.32 0.00	0.13	0.00	MnO	0.16	0.14	0.00	0.12	0.00
CaO 10.11 8.32 0.00	2.40	0.00	MgO	3.14	2.43	0.00	1.74	0.00
	6.90	0.00	CaO	10.63	9.53	0.00	8.60	0.02
Na <sub>2</sub> O 2.12 2.59 0.01	2.64	0.03	$Na_2O$	2.40	2.81	0.00	3.00	0.00
K <sub>2</sub> O 0.70 1.16 0.00	1.67	0.01	$\rm K_2O$	0.58	1.10	0.10	0.90	0.17
$P_2O_5$ 0.16 0.22 0.00	0.25	0.00	$P_2O_5$	0.21	0.26	0.00	0.25	0.01
100.00 100.00	100.00		Sum	100.00	100.00		100.00	
$\sum r^{2}$ 0.13	0.16		$\sum r^2$			0.15		0.48
Removed crystals (wt%)				Re	moved crys	tals (wt%)		
0			0		2.97			
Cpx 6.84	3.38		Cpx		2.58		3.71	
Opx 6.28	4.21		Opx				2.93	
Plg 23.7	16.23		Plg		12.46		17.94	
Fe-Ti oxides 2.86	2.19		Ox		0.35		2.88	
wt% of residual liquid 39.68	73.99		wt% of	residual liqu	iid 81.64		72.54	

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TABLE 9

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Fig. 11 - Plot of CaO (a) and  $Al_2O_3$  (b) vs MgO contents for the investigated Sant'Antioco volcanics, compared with Montresta (data after Morra *et al.*, 1997; rock abbreviations as in Fig. 8). Arrows represent possible differentiation paths, as considered in mass balance calculations (TABLE 8, 9). Samples marked are discussed in the text.

orogenic rocks of Sardinia (Dupuy et al., 1974, 1979; Coulon, 1977; Brotzu et al., 1997a, b; Conte, 1997; Morra et al., 1997; Downes et al., 2001), for which different degrees of crustal contamination, coupled with variable enrichment of mantle source by subducted sediments, has been proposed. At Sant'Antioco, the observed range of Sr isotope data, the positive correlation of 87Sr/86Sr vs elements with crustal affinity (e.g., Si, Rb, Ba; Fig. 8), as well as the wide scattering of trace elements (Fig. 6), strongly suggest contamination by crustal wall rocks. In particular, the Sant'Antioco 87Sr/86Sr values are significantly higher with respect to the chemically analogous products of Montresta, whereas they overlap the higher values from the



Fig. 12 - Representative REE patterns for high-Al<sub>2</sub>O<sub>3</sub> basaltic andesites (HABA) and andesites (AND) from Sant'Antioco and Montresta (shaded field; data after Morra *et al.*, 1997). The pattern for the Montresta high-MgO basaltic sample (HMB, KB13), considered as parental magma in petrogenetic modeling, is also reported (data after Morra *et al.*, 1997).

nearby Sarroch volcanic district (Fig. 8). In the latter, crustal assimilation has been revealed and quantified in about 4-14 wt% (Conte, 1997), where the assimilated material considered, i.e. an average composition of the granitoid batolith of Sardinia (from Del Moro et al., 1975; Secchi et al., 1991), would account for the observed trace element enrichment (e.g., Ba, Rb and Zr) and depletion (e.g., Sr) trends. Besides the oversimplified assumption concerning the crustal material involved, considering the similar <sup>87</sup>Sr/<sup>86</sup>Sr ratios and trace elements behaviour observed for Sant'Antioco and Sarroch volcanics, we infer that crustal contamination affected magma evolution at Sant'Antioco at analogous, or even slightly higher, extent.

# CONCLUDING REMARKS

The present work integrates the petrological knowledge on the Cenozoic orogenic magmatism of western Sardinia, with focus on mafic to intermediate rock-types. New data on Sant'Antioco relatively mafic volcanics, mostly represented by basaltic andesites and andesites, reveal a remarkably wide compositional spectrum in terms of Al<sub>2</sub>O<sub>3</sub>, MgO and CaO contents, as related to the characterizing mineral assemblages. The observed variability in whole rock and mineral chemistry likely correlates to different liquid lines of descent, controlled by the  $P_{H_{2}O}$  histories of magma systems, starting from a common parental magma composition. The latter has been identified as a high-MgO basalt similar in composition to the broadly coeval Montresta primitive lavas. Different from the Montresta district, the Sant'Antioco area is characterized by the lack of primitive terms, possibly indicating different local tectonic regime favouring magma ponding and differentiation at upper crustal levels.

The overall geochemical features of the Sant'Antioco study rock-types evidence a clear calc-alkaline signature, despite relatively high values of FeO\*/MgO that could mimic a misleading tholeiitic affinity. Starting from the composition of a Montresta HMB, considered as a suitable primitive (possibly near-primary) magma, two end-member differentiation trends are recognized, which lead respectively to the MgO-richest/Al<sub>2</sub>O<sub>3</sub>-poorest and MgO-poorest/ Al<sub>2</sub>O<sub>3</sub>-richest basaltic andesites, representing the least evolved rock-types at Sant'Antioco. These trends can be respectively explained by higher proportions of plagioclase vs olivine+Cpx in the fractionating assemblage under H<sub>2</sub>O-poorer conditions and vice versa in H2O-richer conditions. Basically, the  $X_{H_2O}$  imprint is recorded up to the most evolved compositions, being also the main controlling factor of incoming Opx crystallization.

Finally, Sr isotope data and the trends of trace elements of "crustal affinity" with increasing degree of magma evolution, indicate that different degrees of crustal contamination also played a significant role in the petrogenesis of Sant'Antioco volcanics, consistent with magma slow ascent and ponding within a tick crust in multiple reservoirs and/or feeder conduits that behaved independently.

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### APPENDIX A: ANALYTICAL METHODS

Whole-rock samples were analyzed for major and selected trace elements (V, Cr, Co, Ni, Rb, Sr, Y, Zr, Nb, Ba, Th) by a Philips PW1480-XRF spectrometer at IGG-CNR, Rome, using fused dishes (major element) and pressed powder pellets (trace elements). Philips X40 software was used for data reduction. Analytical precision is generally better than 3% and 5%, for major and trace elements, respectively.

LOI was determined after the samples had been kept at 1050°C for 2 h; resulting losses were corrected for FeO oxidation and FeO was measured by titration with KMnO after acid digestion.

REE, Ta, Hf and U were determined by ICP-MS at the Activation Laboratories, Ancaster, Canada.

Sr isotope ratios were measured, after sample dissolution and ion exchange separation procedures, on a VG354 double collector mass spectrometer at the University Federico II, Naples. Repeated analyses of NBS-987 international reference standard gave an average value of  $0.71027\pm0.00001$  (2 $\sigma$ ).

Mineral analyses were carried out using a Cameca SX50 electron microprobe with an accelerating voltage of 15 keV and 15 nA beam current at C.N.R.-IGAG, Rome, using the ZAF correction procedure.

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